# Byurakan Astrophysical Observatory Plate Archive and its scientific usage

Byurakan Astrophysical Observatory (BAO)

Astroplate workshop 18-21 Mar 2014, Prague, Czech Republic

#### **Overview**

**Byurakan Astrophysical Observatory Telescopes and observational material Observational programs accomplished in BAO BAO Plate Archive BAO Electronic Observational Database Digitization projects** Science projects based on BAO Plate Archive data Armenian Virtual Observatory (ArVO), IVOA **Comparison of past and present observations** 

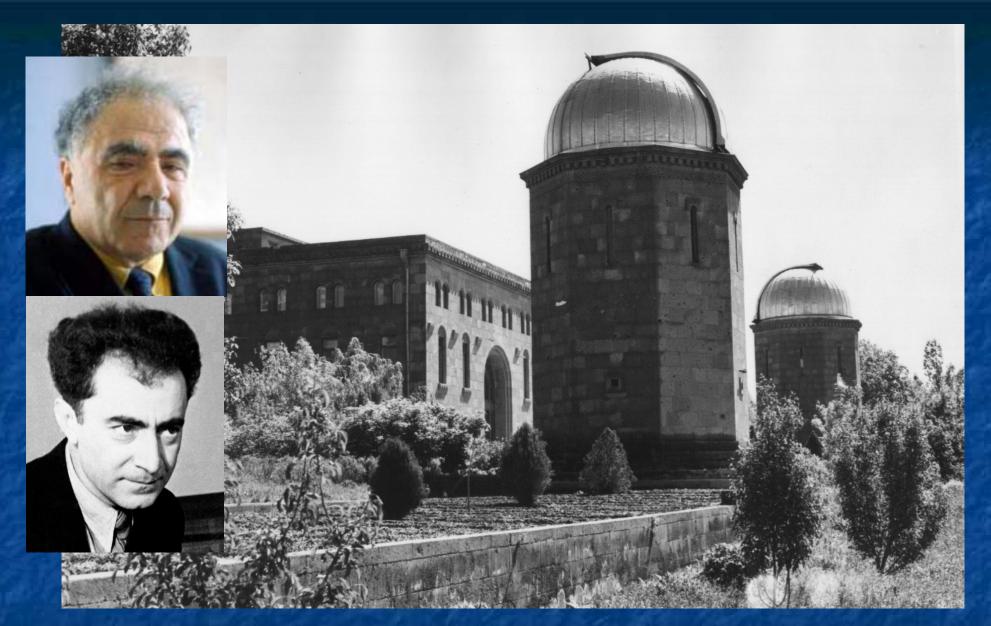
A.M. Mickaelian: BAO Plate Archive and its scientific usage

# Armenia & BAO





Southern slope of Mt. Aragatz, near village Byurakan, 30km Northwest of Yerevan, founded in 1946<sub>3</sub>

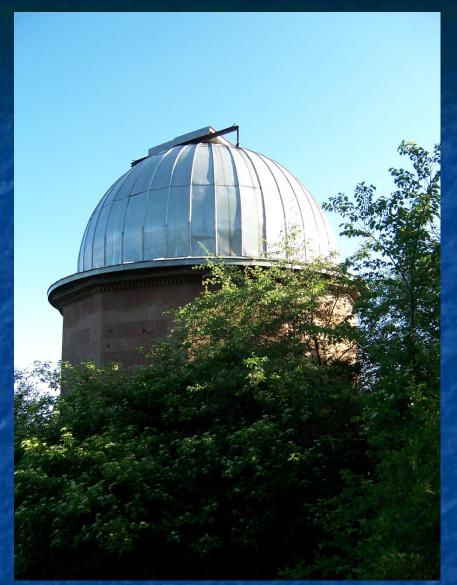


**BAO.** Ambartsumian's discoveries and theories, Markarian, Arakelian, and Kazarian galaxies, Shahbazian compact groups, T Tauri and flare stars, HH objects, SNe, FBS and SBS objects, IRAS sources

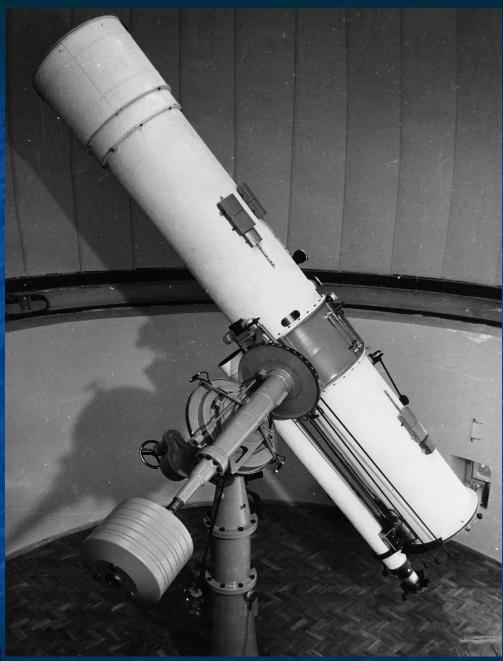
#### **BAO telescopes & observational material**

Telescope name	Size (cm)	Years	Observ. methods	Plates
5" double-astrograph 6" 8" Schmidt 20" Cassegrain 10" telspectrograph nebular spectrograph 16" Cassegrain 21" Schmidt 40" Schmidt (AZT-10) ZTA-2.6m	13 15 20/20/31 51/800 25 41/400 53/53/183 102/132/213 264/1016	1947-1950 1947-1950 1949-1968 1952-1991 1953- 1954- 1955-1991 1955-1991 1960-1991 1975-1991	photometry photometry photometry electrophotometry spectra electropolarimetry photometry photom., spectra photom., spectra	
All telescopes		1947-1991		37000

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Byurakan Observatory 0.5m Schmidt telescope 1955-1991



#### Main observational projects

21" (0.5m) Schmidt telescope:

Polarization of cometary nebula NGC 2261 Nuclei of nearby Sa and Sb galaxies Nuclei of nearby Sc galaxies Search for flare stars in Pleiades Search for flare stars in Orion Search for flare stars in NGC 7000 (Cygnus) Search for flare stars in Praesepe Search for flare stars in Taurus Dark Clouds Variability of Markarian galaxies Monitoring of extragalactic supernovae in certain areas

Byurakan Observatory 1m Schmidt telescope 1960-1991

#### Main observational projects

40" (1m) Schmidt telescope:

Detailed colorimetry of bright galaxies **First Byurakan Survey** (FBS, Markarian survey) Search for flare stars in Pleiades Search for flare stars in Orion Search for flare stars in NGC 7000 (Cygnus) Search for flare stars in Praesepe Search for flare stars in Taurus Dark Clouds **Second Byurakan Survey** (SBS) Extension of the FBS in the Galactic Plane

Byurakan Observatory ZTA-2.6m telescope 1976-1991

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#### Main observational projects

#### ZTA-2.6m telescope:

Morphological study of Markarian galaxies Investigation of star clusters Investigation of groups and clusters of galaxies Spectroscopy FBS blue stellar objects Spectroscopy FBS late-type stars Spectroscopy SBS galaxies and stellar objects (BAO/SAO) Direct images of the central regions of Markarian galaxies Spectroscopy of T Tauri and flare stars Spectroscopy of Byurakan-IRAS Galaxies (BIG objects) Spectroscopy of ROSAT AGN candidates (BAO/HS/OHP/INAOE)



**Byurakan Observatory: other telescopes** 



Yerevan State University (YSU) Observatory 0.2m telescope

#### Library of photographic plates – Plate Archive

Before observers kept plates and films at their offices! There was no rule for storage and preservation.

Plate Archive was organized in 1986, Head: J.A. Stepanian Collected >20,000 plates and thousands of spectra on films; person in charge: G.B. Ohanian (1994-2009) Persons in charge: S.K. Sinamyan (2009-2012), G.M. Paronyan (2012-), members of ArVO group

Digitization team: A.M. Mickaelian, K.S. Gigoyan, L.K. Erastova, L.A. Sargsyan, P.K. Sinamyan

Electronic Plate Archive is now being created >>>

## **BAO Electronic Database Observational projects**

Dates / Julian dates Telescope Observing modes/methods Instrument Receiver Emulsion Filters Seeing **Project name Project PI Observers** 

Targets / coordinates Sky area Surface Scale **Spatial resolution** Spectral range Spectral resolution Limiting magnitude Number of nights Number of exposures Links

## Scanning possibilities

PDS 1010A. Operated accuracy, maximal speed of Stage motion 40 mm/sec, 250x250 mm size
EPSON Expression 16
80 Pro & EPSON Perfection V750 3200 dpi resolution transparency mode (pos-neg), A4 size

Test scanning of DFBS plates with:1995-1996:PDS-1010 (Byurakan)1995-2001:Canon, HP, Epson scanners (Byurakan)1996:APM (Cambridge, UK)1999:MAMA (Paris, France)2002-2003:EPSON Expression 1680 Pro (Rome, Byurakan) \*\*\*2003:StarScan (USNO, Washington, DC, USA)

## The First Byurakan Survey (FBS)

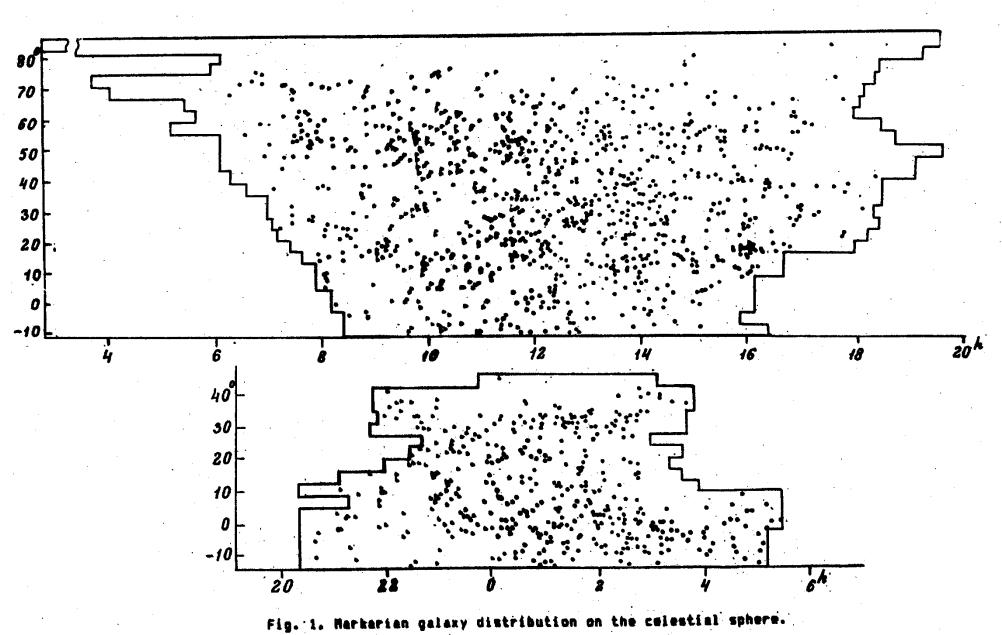
Authors: Years: Instruments:

Emulsions: Spectral range: Field: Scale: Region of sky: Total area:

Limiting magn: Main goal: Selected objects:

Number of objs: Publication:

B.E.Markarian, V.A.Lipovetsky, J.A.Stepanian 1965-1980 102/132/213 cm Byurakan Schmidt telescope 1.5° objective prism (1800 Å/mm at  $H_{\gamma}$ ) Kodak IIA-F, IIa-F, 103a-F, II-F 3400-6900 Å with a sensitivity gap near 5300 Å 4.1° × 4.1° (plates: 16 × 16 cm) 96.8 "/mm  $\delta \ge -15^{\circ}, \ |b| > 15^{\circ}$ 17,056 sq. degree (1139 fields, more than 1874 plates)  $17^{m}$ -17.5<sup>m</sup> ( $\leq 18.5^{m}$ ) selection of UV-excess galaxies UV gal, Sy, QSO, BLL, LINER, SB, HII, radiogalaxies, etc. 1515 15 lists (1967-1981), 2 catalogs (Mazzarella & Balzano 1986, Markarian et al. 1989)



## **The Markarian Survey**

- First systematic objective-prism survey
- The largest objective-prism survey of the Northern sky (17,000 sq. deg)
- New method of search for AGNs
- 1515 UVX galaxies: 181 Seyferts, 17 LINERs, 13 QSOs, 3 BLLs, 95 Starburst, 26 HII galaxies
- Classification of Seyferts: Sy1 & Sy2
- Definition of Starburst galaxies
- Other projects: FBS BSOs, RS, opt. ident. (BIG & BIS objects)

# Digitized DDFBS First Byurakan Survey

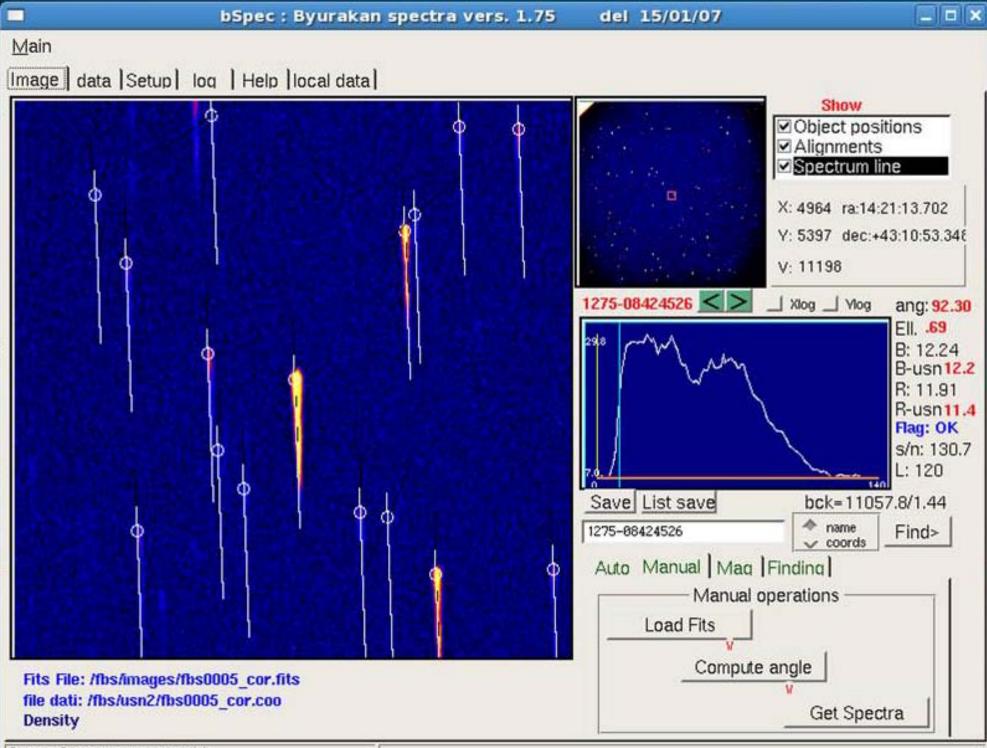
Collaboration between Byurakan Astrophys. Observ. (BAO, Armenia), La Sapienza Univ. di Roma (Italy), and Cornell Univ. (Ithaca, NY, USA)

scanning astrometric solution extraction wavelength calibration density and flux calibration multiband (UBVR) photometry making up template spectra numerical classification **-**17 DFBS catalog and database web page and user interface

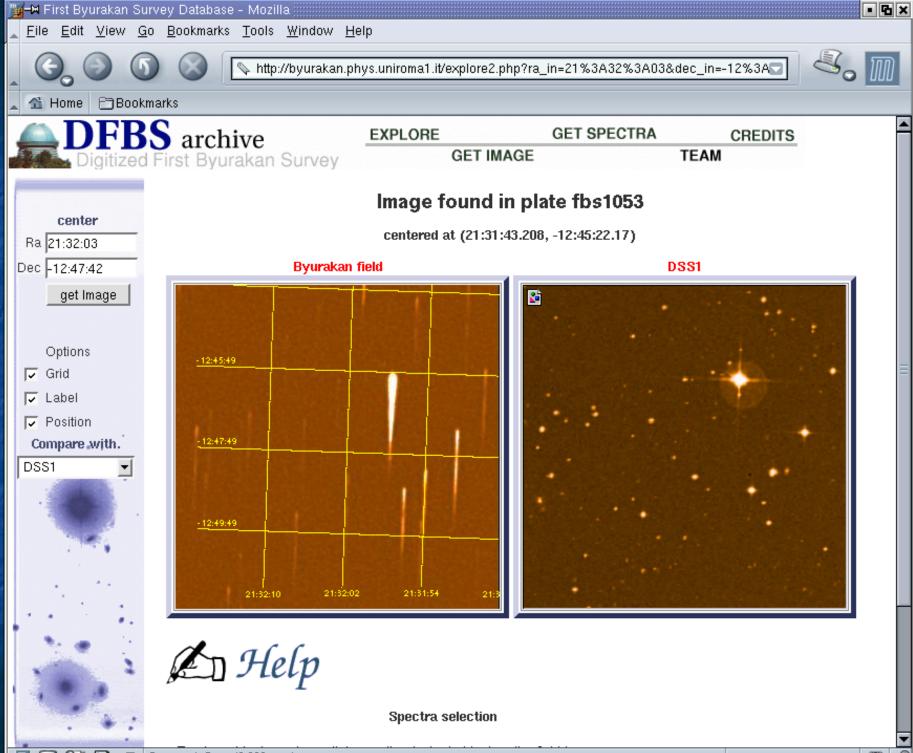
10, 05 Aug 2002 15:41:25	MIDAS version: 01FEB
156	1843

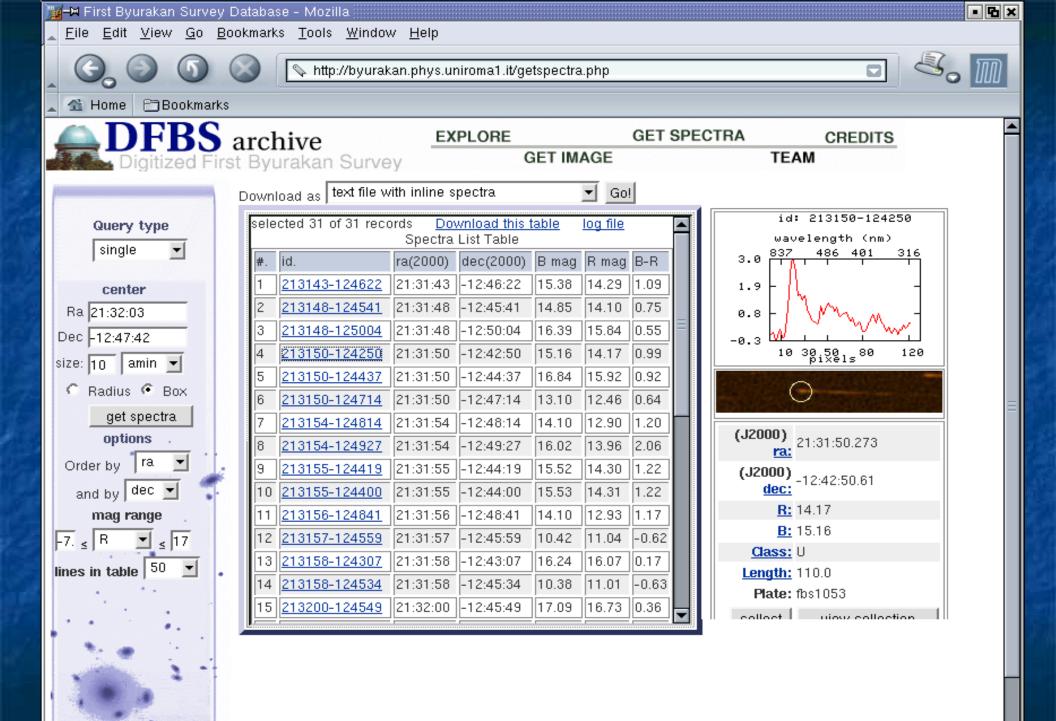
#### Main features:

Pixel size  $15.875\mu$  or 1.542" 9601×9601 pixels each plate 180 MB file for each plate 107*pix* × 5*pix* spectra  $(1700\mu \text{ length})$ Astrometric solution ~1" rms Dispersion: 33A/pix (22-60 from B to R) **Photometric** accuracy ~0.3m 1874 plates, ~400GB 40,000,000 spectra for 20,000,000 objects<sub>21</sub>



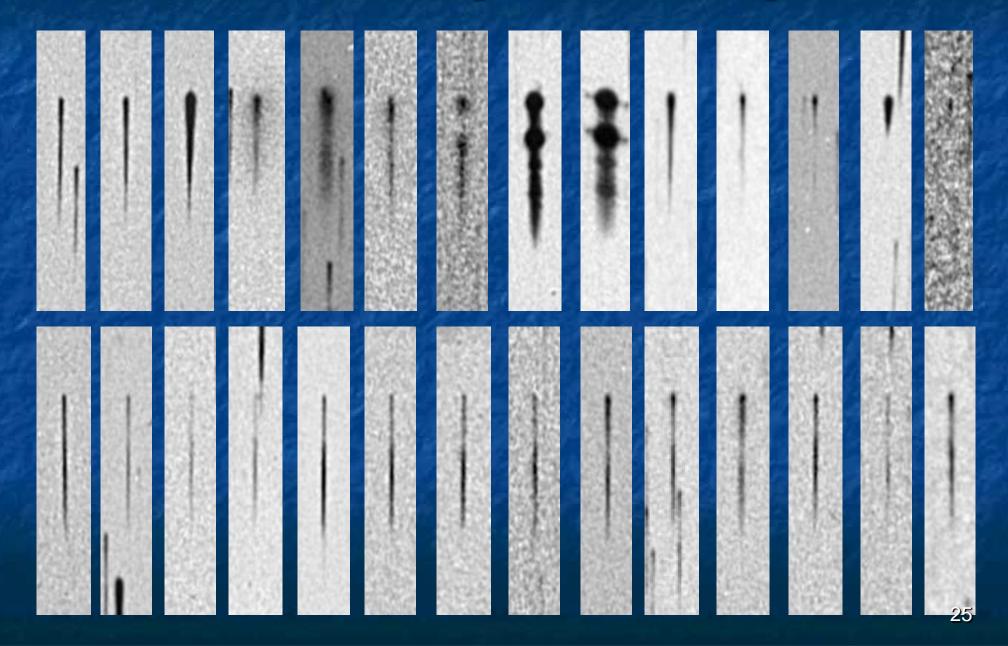
Spec: Oggetto n.10734 id=



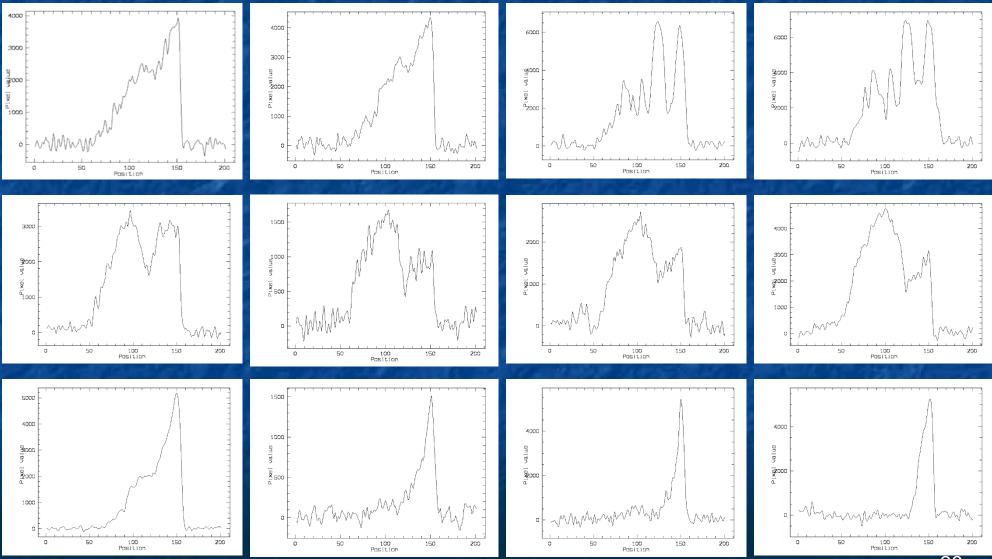


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# **DFBS low dispersion spectra**



# **DFBS low dispersion spectra**



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#### Other digitization projects: Second Byurakan Survey (SBS)

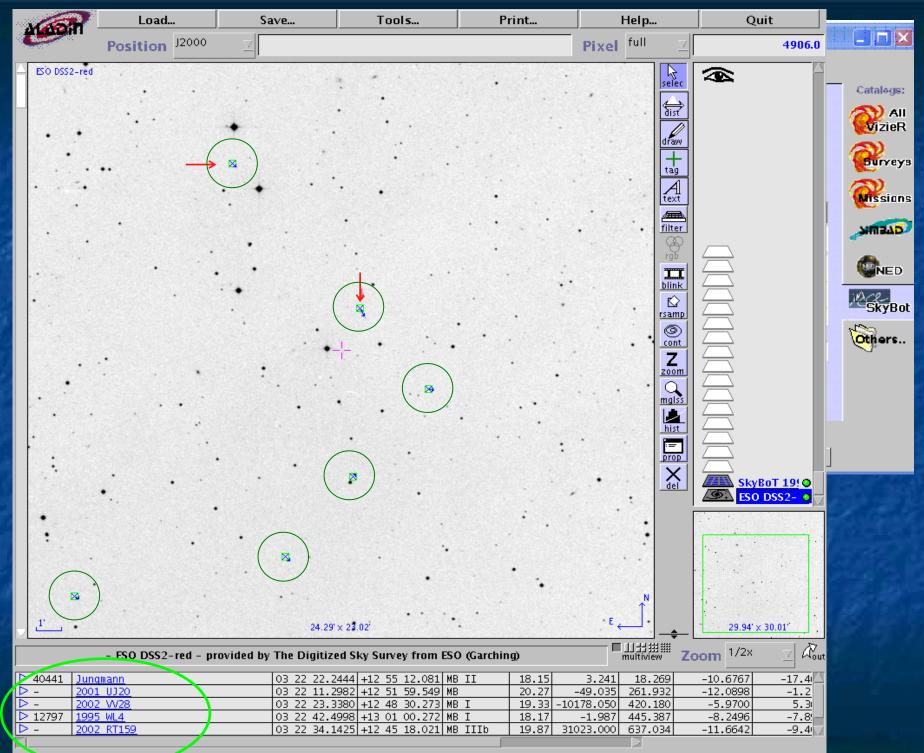
B.E.Markarian, J.A.Stepanian, L.K.Erastova, V.H.Chavushyan **Authors:** 1978-1991 Years: **Instruments:** 102/132/213 cm Byurakan Schmidt, 1.5°, 3° & 4° prisms (1800 Å/mm, 900 Å/mm & 280 Å/mm at Ηγ) Baked Kodak IIIa-J, IIIa-J+GG495, IIIa-F+RG2, IV-N **Emulsions:** Spectral range: 3400-5300 Å, 4950-5400 Å, 6300-6950 Å Field: 4.1°× 4.1° (plates: 16×16 cm) Scale: 96.8 "/mm **Region of sky:**  $49^{\circ} \le \delta \le 61^{\circ}$ ,  $|b| > 30^{\circ} (7^{h} 4 3^{m} \le \alpha \le 17^{h} 15^{m})$ Total area: *965 deg<sup>2</sup> (65 fields, 550 plates)* **Limiting magn:**  $18^m$ - $20^m$  in V (completeness is  $\leq 17.5^m$ ) Main goal: Extension of the FBS to fainter magnitude limits Methods: UVX / emission lines / SED **Digitization:** *since 2003: 16 bit, 2400 dpi (10µ pixel size); 180 SBS plates* 

# Science projects based on BAO Plate Archive data

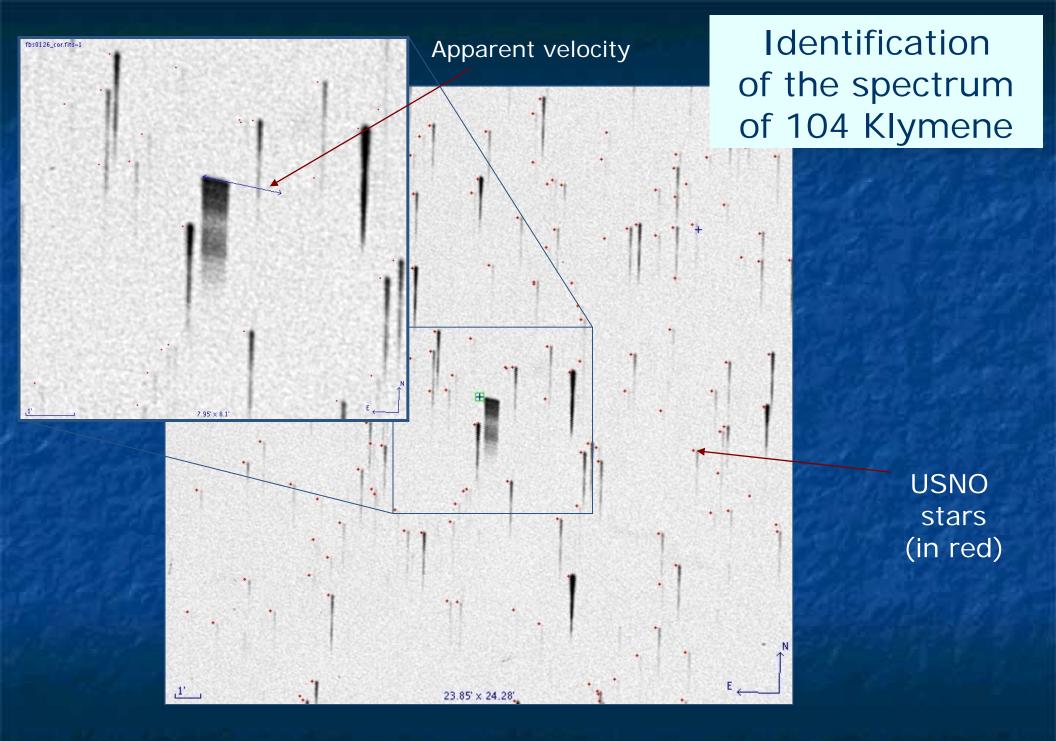
## **Search for Asteroids in the DFBS**

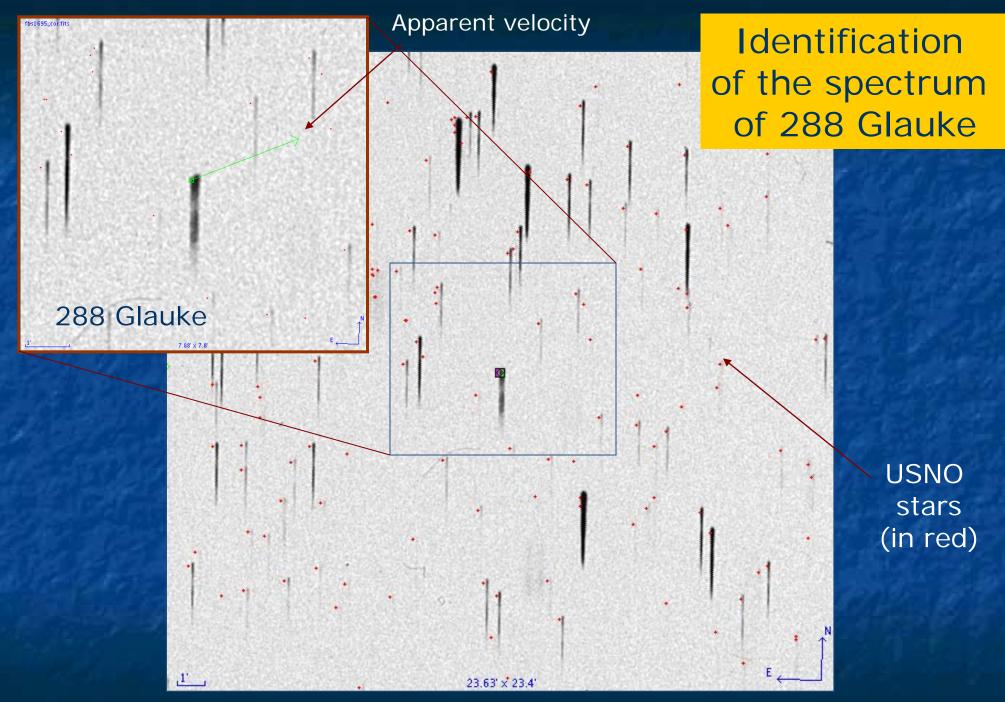
#### Aladin and SKYBOTE

- the brightest asteroids (<15-16), which may be visible in the DFBS plates</p>
- fast & slow asteroids with a division parameter, estimated as the motion of 3" during 20 minutes (the typical exposure time of a DFBS plate)
- extraction of all spectra of asteroids found in DFBS by SkyBote; group them into extended (fast asteroids) and starlike (slow asteroids)
- modeling of a template spectra of asteroids by means of the star-like spectra
- search for new candidate asteroids by similar spectra & comparison with DSS1/DSS2 fields for elimination of the stars
- spectral analysis of the asteroid spectra to get some physical parameters



(c)1999-2005 ULP/CNRS - Centre de Données astronomiques de Strasbourg

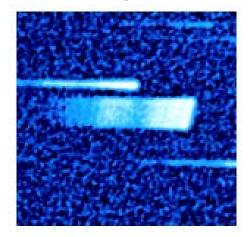




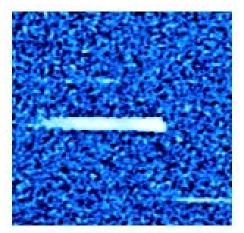
مغارب والتركي والمتعادين

#### **Spectra of asteroids in the DFBS**

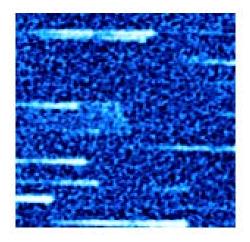
#### 104 Klymene



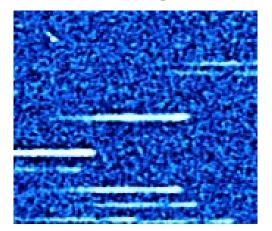
288 Glauke



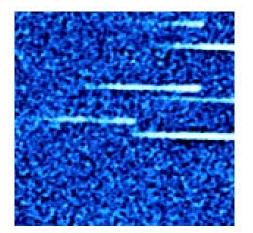
627 Charis



1323 Tugela



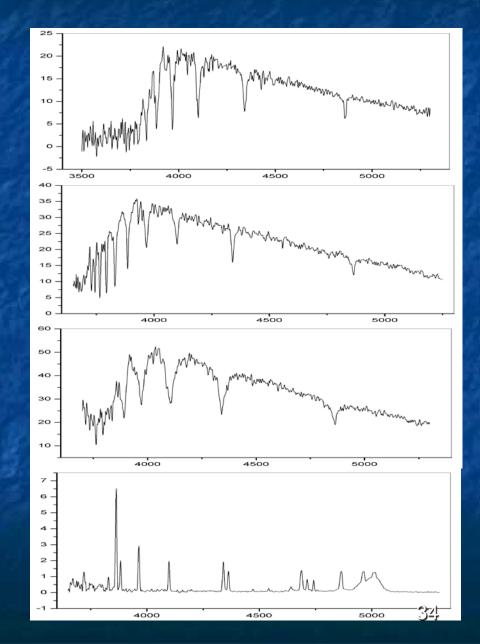
**4713 Steel** 

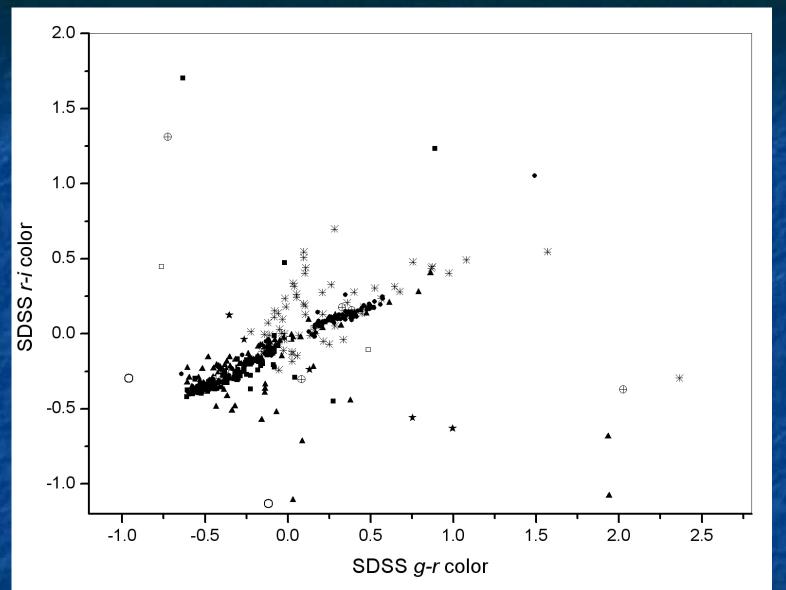


## **Spectroscopic study of the FBS BSOs**

Scanning spectra: 1600 dpi, 16 bit, 650x21 pix sizes ~700 spectra have been digitized (FBS blue stellar objects & late-type stars), 101 published

A new PN (FBS 2323+421), 7 white dwarfs (DA, DAB, DZ), 78 hot subdwarfs, 9 HBB stars have been revealed.





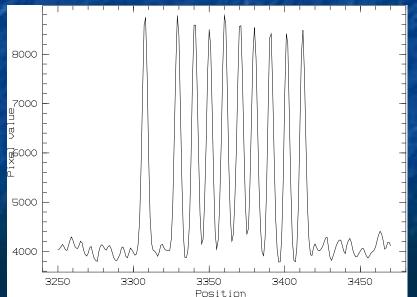
Asterisks – extragalactic objects (AGN and Gal) Open circles – **PNN** Stars – CV Filled squares – WD Triangles – hot subdwarfs (sd)Open squares – HBB Earth symbols – MS stars Filled circles – unknown objects

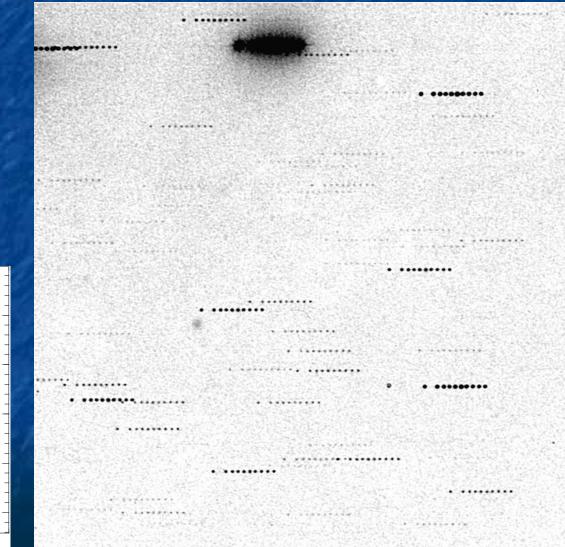
Sinamyan P.K., Mickaelian A.M. Astrophysics **49**, 333 (2006); **51**, 51 (2008); **51**, 226 (2008); **52**, 76 (2009)

### **Study of long-term variability of ON 231**

Photographic chains for discovery of flare stars in stellar clusters

Coma: 189 plates with a total number of more than 1200 exposures in 1965-1976 with the Byurakan 21" and 40" Schmidt telescopes



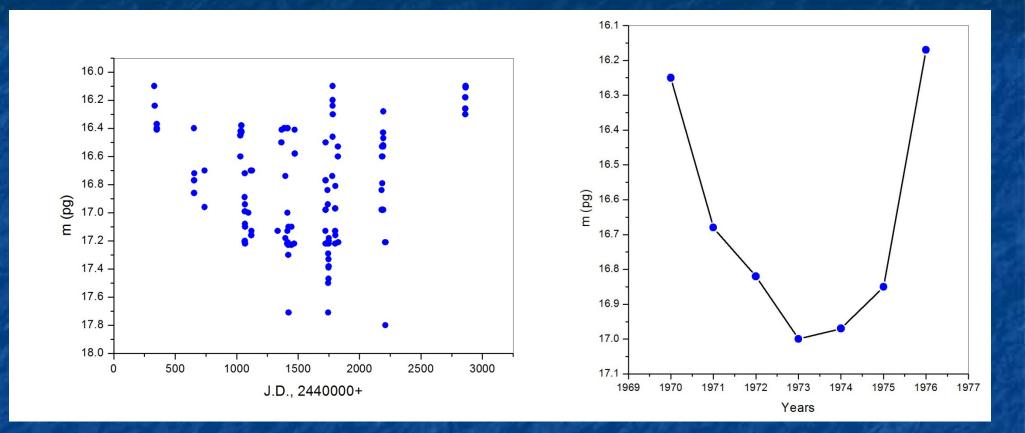


#### **Study of long-term variability of ON 231**

#### **ON 231** (W Com, m=16.5, z=0.102)

- 1969-1976, L.K. Erastova, observations in Coma field (R.A.=12<sup>h</sup>26<sup>m</sup>, DEC=+25°55') for study of the variations of the flare stars
- 187 plates with "chain" images (by a slight shifting of the plate for multiple images), BAO 0.5m and 1m Schmidt telescopes (5°×5° and 4°×4° fields, respectively)
- ON 231 was present in 141 of these plates with some 1200 total exposures All available plates were scanned
- Out of 141 plates with ON 231, useful measurements were done in about 80 ones
- A script was written under MIDAS for reduction of the plates and measurements
- The study of the variability was done by a comparison with the neighbouring standard stars, and the equivalent widths (EW) of the measured images (each point in the chains) were the main criteria for the comparison

## **Study of long-term variability of ON 231**



**Two other AGN in the field** (will be studied later for possible variability): FBQS J122424.3+243623 (z=0.218, a 87GB and ROSAT source) FBQS J123014.1+251804 (z=0.135, a ROSAT source)

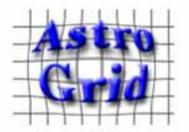
### **Other papers at Astroplate 2014**

**Roberto Nesci:** AGB stars and the plate archives heritage

Variability of IRAS-selected AGB stars identified with the objective-prism plates of the FBS

Rene Hudec: Low Dispersion Spectroscopy (LDS) with astronomical photographic databases and surveys Applications in modern astrophysics (e.g. ESA Gaia project)

> Armenian Plate Archive and Digitization Projects















VIRTUAL ASTRONOMICAL OR





SA<sup>3</sup>

South African Foundation Astroinformatics Alliance

NRF

National Research













Hungarian Virtual Observatory





## **BAO National Value project on Plate Archive**



Creation of the project team: astronomers, technicians, computer scientists **Renovation of BAO Plate Archive area** Creation of BAO full electronic database; interactive electronic journals, etc. Purchase of new scanning equipment Digitization of plates by importance of scientific projects; ~ 10,000 plates during 2014-2015 Creation of interactive sky map for all BAO observations New science projects based on BAO Plate Archive

### **Past and present observations**

#### **Numbers**

Telescopes Wavelength coverage Limiting magnitude Spatial/spectral resolution Total observing time Number of exposures Number of observed objects Number of bytes Time coverage Sky coverage Number of results Number of finances

hotographic	Digital
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×	×
0	×
<b>619</b> × 0 0	×
0	×
×	×
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×	×
× ×	×
×	<b>0</b> 42
	42